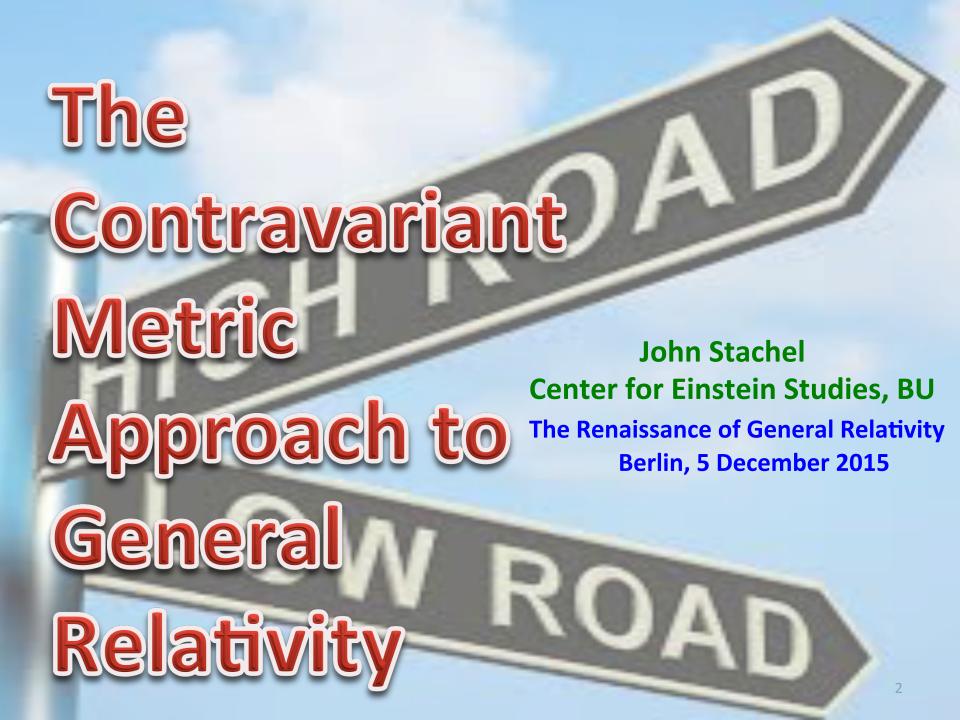
# The Contravariant Metric Approach to General Relativity

John Stachel
Center for Einstein Studies, BU
The Renaissance of General Relativity
Berlin, 5 December 2015



#### **Contravariant Metric**

We shall take the contravariant metric  $g^{\mu\nu}$  and covariant vectors  $\omega_{\mu}$  as the starting point for discussion of the pseudo-metric. Traditionally, of course, just the opposite is done: one starts with the covariant metric  $g_{\mu\nu}$  and contravariant vectors  $V^{\mu}$ .

But the physical interpretation of the firstorder formalism for general relativity suggests taking the contravariant approach.

## First-Order Lagrangian

The Lagrangian density is:

$$\mathcal{L} = (-g_b)^{1/2} g^{\mu\nu} A_{\mu\nu} (\Gamma, \partial \Gamma)$$
$$= (-g^{\#})^{-1/2} g^{\mu\nu} A_{\mu\nu} (\Gamma, \partial \Gamma).$$

In this equation,  $A_{\mu\nu}$  ( $\Gamma$ ,  $\partial\Gamma$ ) is the affine Ricci tensor written as a function of the symmetric affine connection  $\Gamma$  and its first derivatives  $\partial\Gamma$ 

Thus, only the contravariant metric is needed to derive the gravitational field equations, as we shall discuss in detail later.

# Take the high road the view is better

## So We'll Take the High Road!



# My Apologies in Advance

Time limits require brevity and brevity is the mother of dogmatism.

None of my statements should be interpreted dogmatically—they are meant to stimulate critical thinking and further discussion.

#### "Flat" versus "Sharp"

Quantities formed from the covariant metric  $g_{\mu\nu}$  will be called "flat" & distinguished by the symbol ","; while quantities formed from the contravariant metric  $g^{\mu\nu}$  will be called "sharp" & distinguished by the symbol "#".

#### **Space-Time Structures in GR**

Zeroth order: 4-Dim Differentiable Manifold  $X_4$ Diff  $(X_4)$ , Unimodular SDiff $(X_4)$ , Tangent and Cotangent Spaces, Volume elements

First Order: Affine Connections, Pseudo-metrics and Connections, Compatibility conditions Conformal and Projective structures, Torsion

Second order: Affine, Metric, Conformal and Projective Curvature tensors and a Ricci tensor for each curvature tensor

#### **Space-Time Structures in GR**

I shall emphasize certain relations between these structures:

**Compatibility between Two Co-Determination of a Third** 

NB: All structures are four-dimensional unless otherwise indicated

```
(e.g., "volume" = "four-volume")

"Metric" = "pseudo-metric with Lorentz signature (+ - - -)"
```

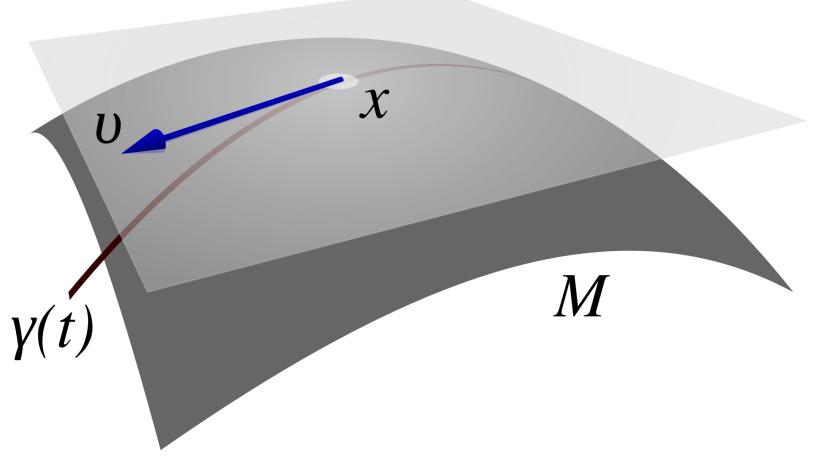
#### Differentiable Manifold: Zero<sup>th</sup> Order Structures

4-Dim Differentiable Manifold  $X_4$  with Symmetry Group:  $Diff(X_4)$ 

Tangent Space  $T_p(X_4)$  at each  $p \in X_4$ Centered Affine Space with Symmetry Group: GL(4, R)

### **Tangent Space**

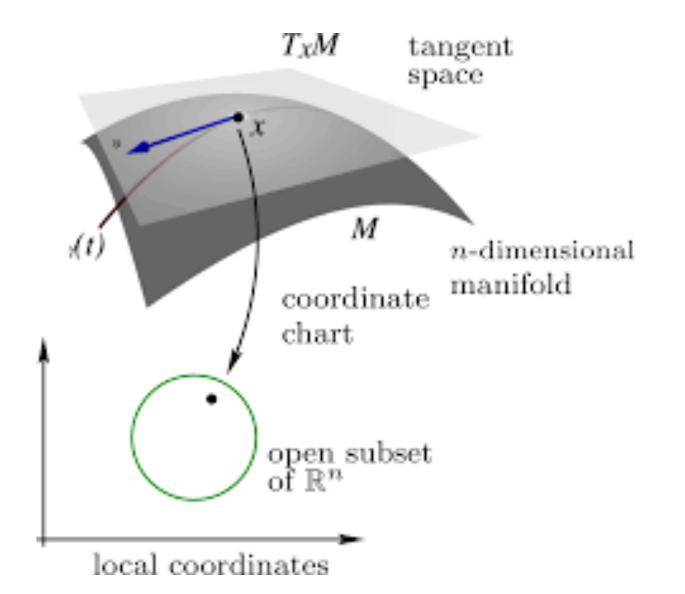
 $T_{\chi}M$ 



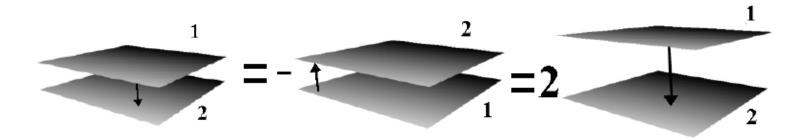
## The General Linear Group and its Subgroups

- $G_a$  The affine group  $\Delta \neq 0$ ,  $E_n$  Affine Space
- $G_{ho}$  Homogeneous trans'ns  $\Delta \neq 0$ , Centered  $E_n$
- $G_{eq}$  Equivoluminar  $\Delta = \pm 1$ ,  $E_n$  with given unit volume
- $G_{ea}$  Special affine trans'ns  $\Delta = +1$ ,  $E_n$  with given unit volume and screw sense
- $G_{pso}$  Pseudo-orthogonal trans'ns  $\Delta \neq 0$ ,  $C_n$  (rotations, reflections and similarity transformations)
- $G_{pson}$  Pseudorthonormal trans'ns  $\Delta = \pm 1$ ,  $R_n$  (rotations and reflections)
- $G_{psor}$  Pseudo-rotations  $\Delta = 1$ , Oriented  $R_n$  (rotations)

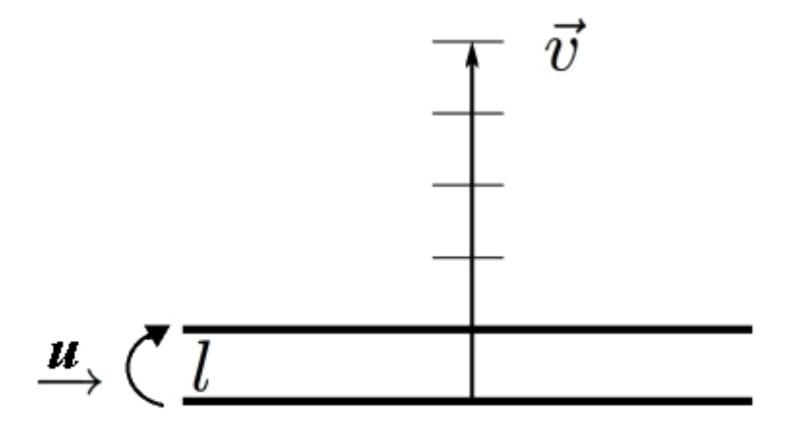
#### Coordinatization



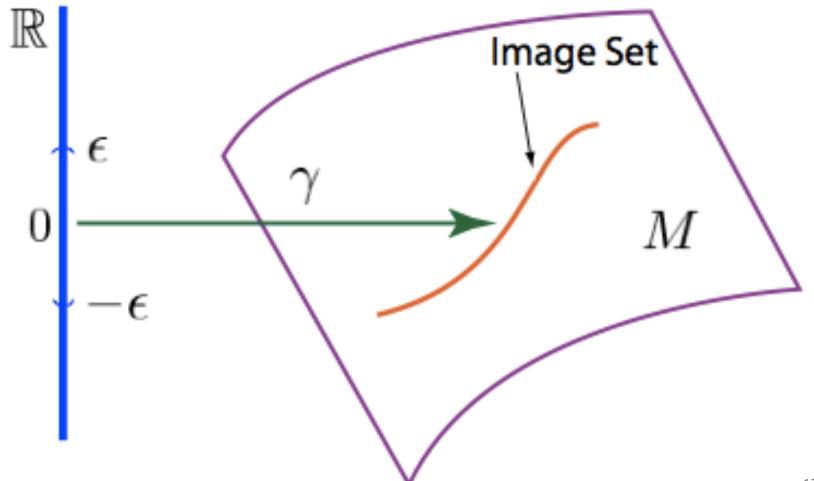
#### **Vectors and Covectors**



# Contraction of vector and covector



## A Curve on the Manifold M is a smooth map $\gamma$ : $-\varepsilon < R < \varepsilon \rightarrow M$



#### Zeroth Order Dualities

#### Differentiable Manifolds

Vectors **V** 

Covectors w

Dual Pair if  $(\omega, V) = 1$ 

Multivectors V [...]

Forms  $\omega$  [...]

 $T(X_n)$  tangent space  $T^*(X_n)$  cotangent space

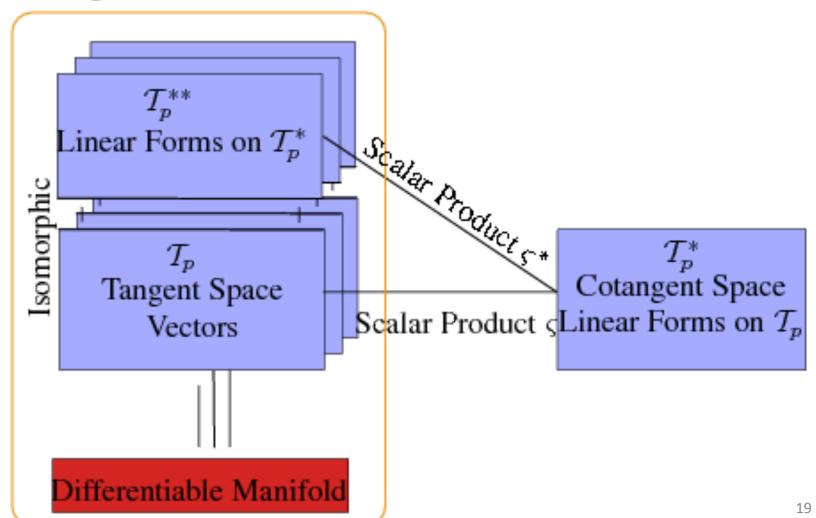
**Fibration** 

**Foliation** 

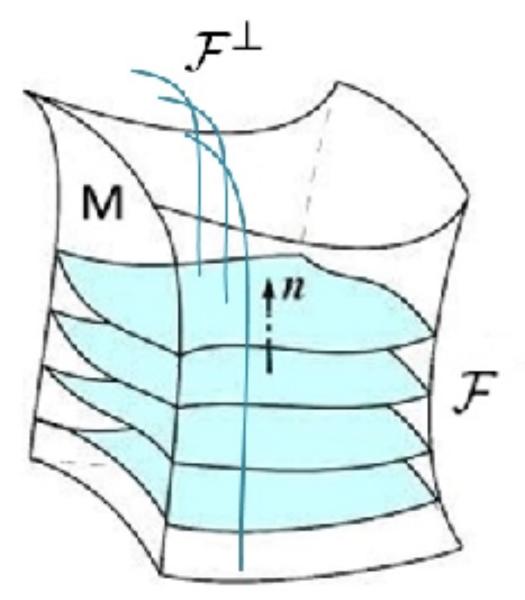
#### **Dual Spaces:**

#### Tangent $T(X_n)$ , Cotangent $T^*(X_n)$

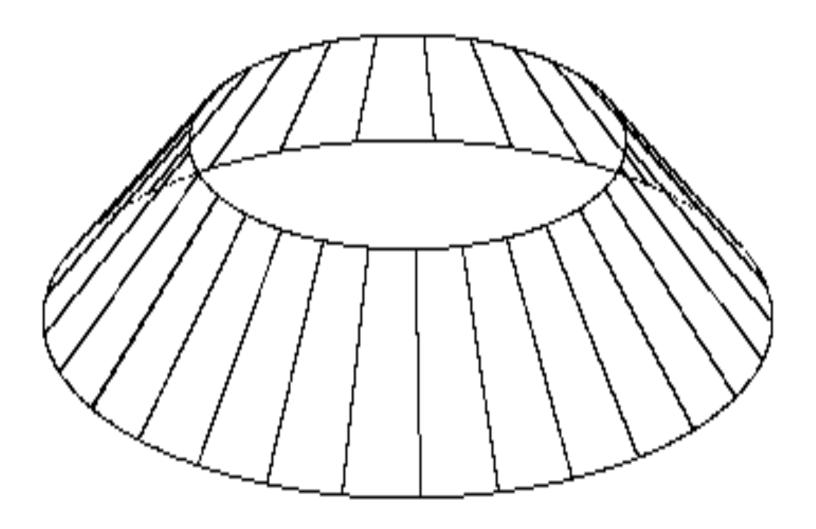
Tangent Bundle



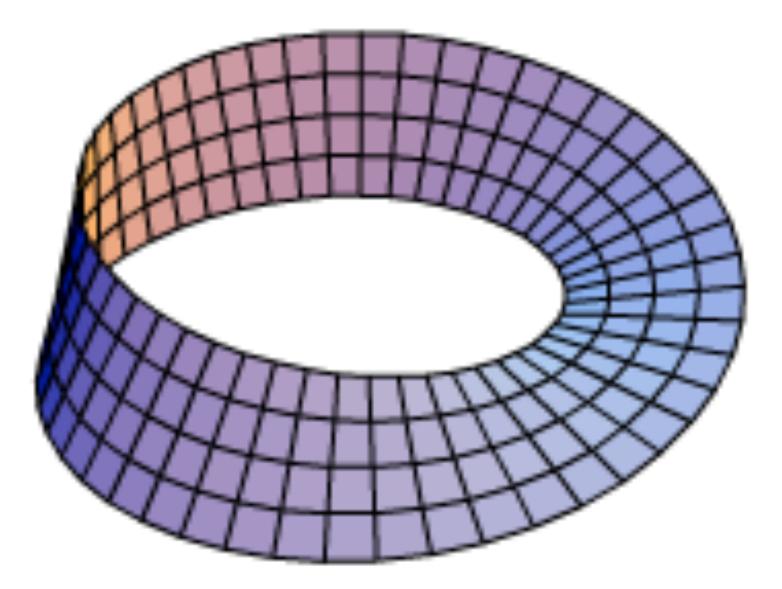
## Foliation



### **Fibration**

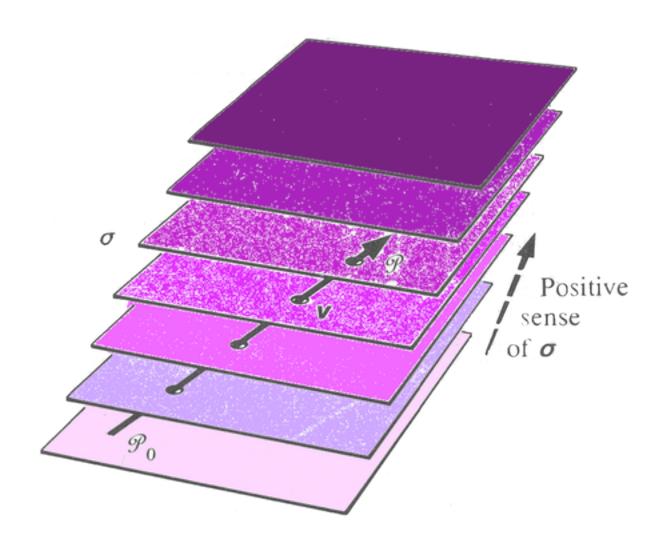


#### **Fibration and Foliation**

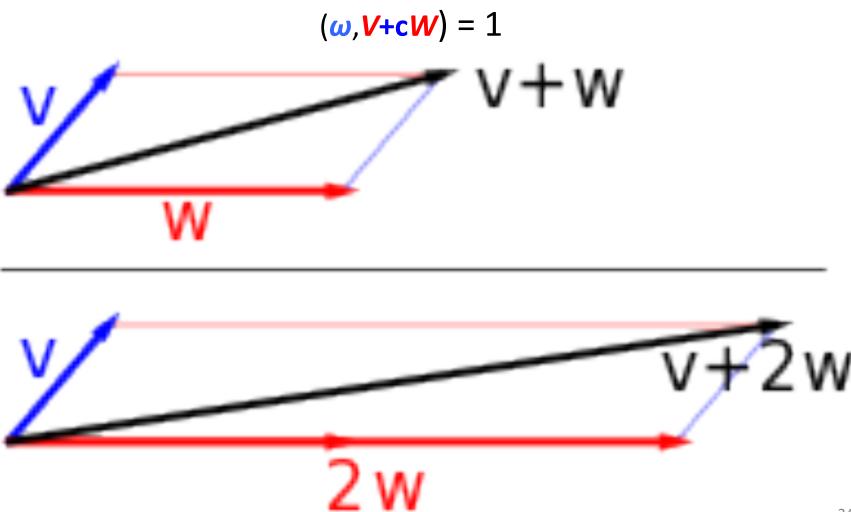


#### **Foliation and Dual Fibration**

 $(\boldsymbol{\omega}, \boldsymbol{V}) = 1$ ,  $\boldsymbol{V} = d\boldsymbol{x}/d\sigma$ 



## Invariantce of Duality Under Certain Changes of V:



#### Some Things Einstein Got Right!!

1) The Hole Argument: Points of Space-Time have no inherent physical properties-- They inherit *all* of these properties from the Space-Time Structures, including all fields

Conclusion: No first order space-time structures or fields, no space-time!

If the water drains out of this bathtub, it takes the bathtub with it!

## GR is a Background-Independent Theory

The contrast between general relativity and all previous theories:

Background-dependent theory: Fixed and given space-time stage, on which the drama of physics unfolds.

Background-independent theory: No actors, no stage, no anything.

Einstein put it this way:

"Space-time does not claim existence on its own, but only as a structural quality of the field."

## "Relativity and the Problem of Space" (1952)

On the basis of the general theory of relativity ... space as opposed to 'what fills space' ... has no separate existence. If we imagine the gravitational field to be removed, there does not remain a space of the type [of SR], but absolutely nothing, not even a 'topological space'.

#### Realizing Einstein's Vision

The concepts of fiber bundles and sheaves enable a mathematical formulation of general relativity consistent with Einstein's vision.

But we have no time to go into this. See, e.g., the PowerPoint

John Stachel
Center for Einstein Studies
Boston University
Florence Category Day
16 June 2010

#### And the paper:

"The Hole Argument and Some Physical and Philosophical Implications,"

Living Reviews in Relativity, 2014

## Differentiable Manifold: First Order Structures

Pseudo-Metric with Lorentz signature (+ ---), Minkowski Space is the tangent space with symmetry group O(3,1)

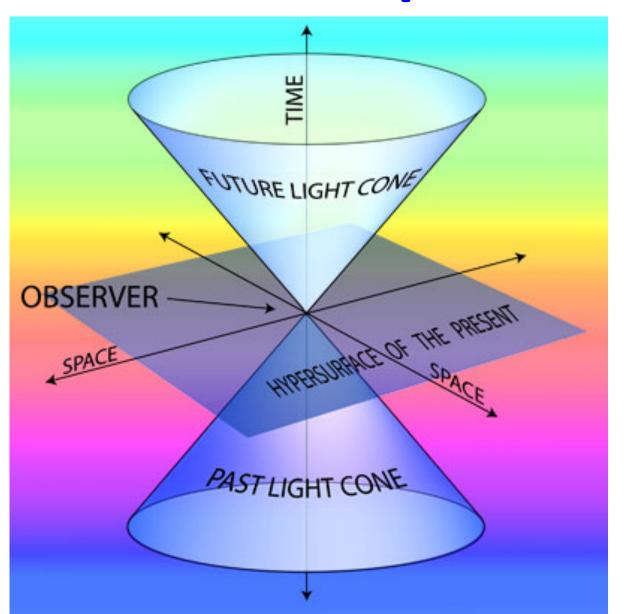
Affine Linear Connection, covariant derivative and affine curvature tensor with symmetry group Aff(4)

**Compatibility Conditions** 

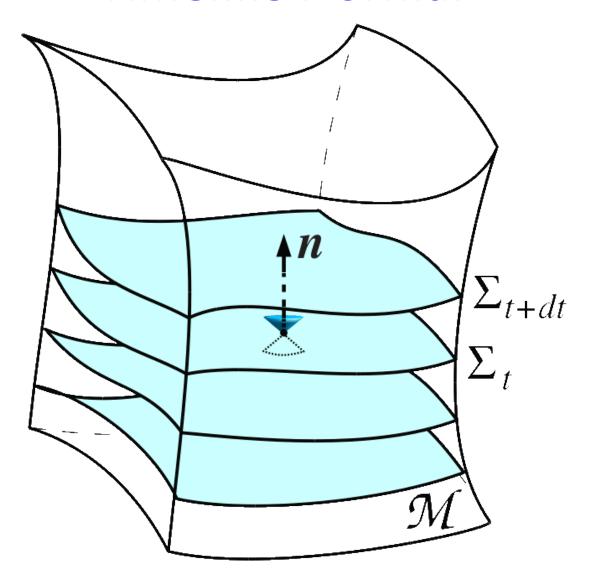
## Reasons for Adopting the Equivoluminar Condition $\Delta = \pm 1$

- 1) To restrict the pseudo-orthogonal group locally and the pseudo-orthogonal metrics globally, resulting in the invariant decomposition of the pseudo-orthogonal metric into an equivoluminar (det = -1) pseudo-orthogonal metric and a scalar field.
- 2) To restrict the affine geometry locally and the linear affine connection globally, resulting in the invariant decomposition of the connection into a trace free connection and a one form.

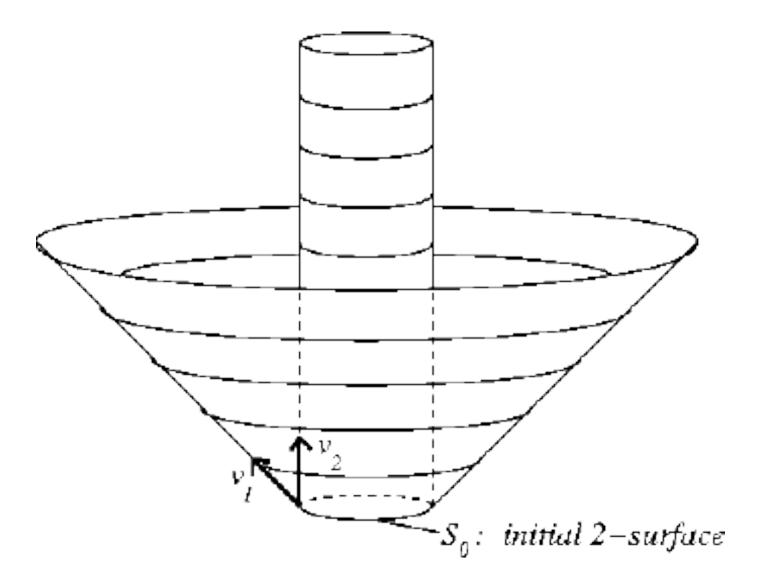
#### SR: Minkowski Space-Time



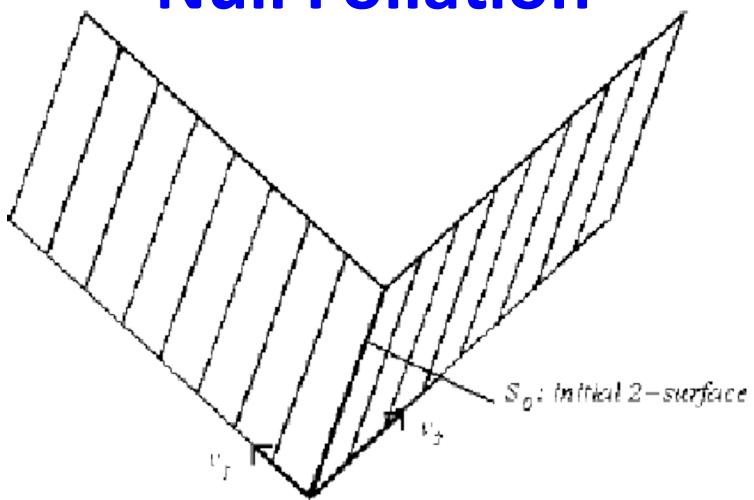
## Lorentz Metric: Spacelike Foliation, Timelike Normal



#### **Lorentz Metric: Null Hypersurface**



# Lorentz Metric: Double Null Foliation



#### **Conformal Two-Structure**

An analysis of the field equations shows that the two gravitational degrees of freedom may be chosen, with no constraints, as the conformal 2- structure.

The dynamical equations then consist of two equations, which propagate the conformal 2-structure.

See R.A. d'Inverno and J. Stachel, J. Math, Phys. 19 (1978): 2447.

#### First Order Dualities

#### **Pseudo-Metric**

```
g_b(V,V) g^\#(\omega,\omega)

V is timelike if g_b(V,V) > 0 \omega is spacelike if g^\#(\omega,\omega) > 0

" " null " " = 0 " " null " " = 0

" " spacelike " " < 0 " " timelike " < 0
```

#### **Metric Connection**

$$D_b \qquad D^{\#}$$

$$\{m_{ab}\}_b \qquad \{nm_a\}^{\#}$$

# Reminder: "Flat" versus "Sharp"

Quantities formed from the covariant metric  $g_{\mu\nu}$  will be called "flat" and distinguished by the symbol "h"; while quantities formed from the contravariant metric  $g^{\mu\nu}$  will be called "sharp" and distinguished by the symbol "#".

# "Flat" Christoffel Symbols

The "flat" Christoffel symbols of the first kind, are defined by:

$$[\lambda v, \mu]_b = \frac{1}{2} (-\partial_{\mu} g_{\lambda v} + \partial_{\nu} g_{\mu \lambda} + \partial_{\lambda} g_{\nu \mu}).$$

The usual "flat" metric connection  $\binom{\kappa}{\lambda\nu}$  is defined as:

$$\{\kappa_{\lambda\nu}\} = [\lambda\nu,\mu]_b g^{\kappa\mu}$$
.

It follows that:

$$D_{\mu} g_{\lambda \nu} = \partial_{\mu} g_{\lambda \nu} - [\mu \lambda, \nu]_b - [\mu \nu, \lambda]_b$$

# "Sharp" Christoffel Symbols

The "sharp" Christoffel symbols of the first kind, are defined by:

$$[\lambda v, \mu]^{\#} = \frac{1}{2} (g^{\kappa\mu} \partial_{k} g^{\lambda v} - g^{\kappa v} \partial_{k} g^{\mu \lambda} - g^{\kappa \lambda} \partial_{k} g^{\nu \mu}).$$

The usual "flat" metric connection  $\{\mu_{\alpha\beta}\}$  can now be defined as:

$$\{\mu_{\alpha\beta}\} = [\lambda \nu, \mu]^{\#} g_{\nu\alpha} g_{\lambda\beta}$$

# "Sharp" Christoffel Symbols

But we now define "sharp" Christoffel symbols of the second kind:

$$\{v_{\alpha}^{\mu}\}^{\#}=_{\mathrm{df}}[\lambda v,\mu]^{\#}g_{\lambda\alpha}$$

from which it easily follows that:

$$D_{\alpha}g^{\lambda\nu}=\partial_{\alpha}g^{ln}+\{\nu\lambda_{\alpha}\}^{\#}+\{\lambda\nu_{\alpha}\}^{\#}=0.$$

We shall see that there is no need for the "flat" Christoffel symbols.

# "Sharp" Covariant Derivatives

It follows from this definition that:

$$D^{\mu}g^{\lambda\nu} = g^{\kappa\mu}\partial_{k}g^{\lambda\nu} + [\mu\nu,\lambda]^{\#} + [\mu\lambda,\nu]^{\#}.$$

Similarly, by using the sharp Christoffel symbols  $\{v_{\alpha}^{\mu}\}^{\#}$ , we can show that

$$D^{\mu}V^{\nu} = g^{\kappa\mu}\partial_{\kappa}V^{\nu} + \{\mu\nu_{\alpha}\}^{\#}V^{\alpha}$$

and

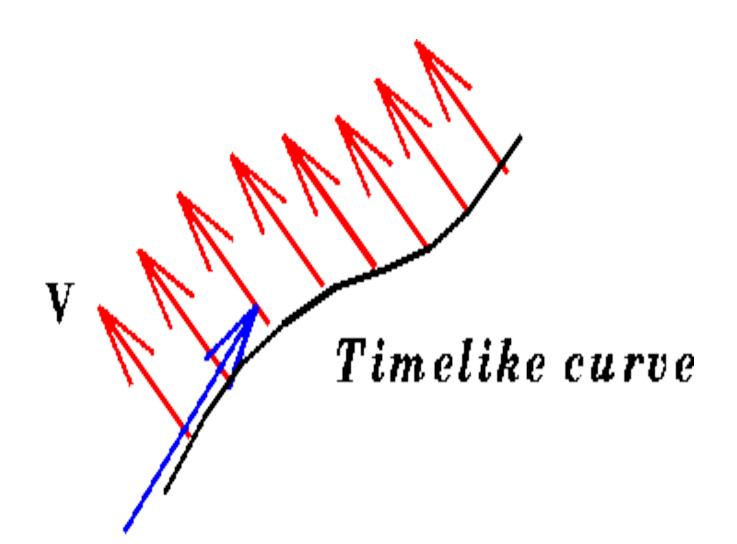
$$D^{\mu}\omega_{\nu} = g^{\kappa\mu}\partial_{\kappa}\omega_{\nu} - \{^{\mu\lambda}_{\nu}\}^{\#}\omega_{\lambda}.$$

# The Inertial Structure

This determines the motion of freely-falling (i.e., net force-free) bodies ("particles") in both classical and special-relativistic physics:

They follow the straightest time-like, inertial paths (autoparallels) of space-time-- straight lines for the "flat" (curvature = 0) space-times of both N-G and Minkowski space-time.

# Parallel Transport Along a Time-Like Curve



# **Autoparallel Curves**

A curve is a parametrized path.

If the tangent vector field to a curve is parallel transported into itself along the curve, it is called an autoparallel curve (& often loosely called "a geodesic")

#### **Affine Connection**

The affine connection  $\Gamma^{\kappa}_{\mu\nu}$  represents the inertio-gravitational structure of space-time.

The affine connection  $\Gamma^{\kappa}_{\mu\nu}$  determines the autoparallel behavior of free particles

We assume the connection is symmetric:

The torsion tensor 
$$S^{\kappa}_{[\mu\nu]} = \Gamma^{\kappa}_{[\mu\nu]} = 0$$
.

#### Law of Inertia and Autoparallels

The affine connection  $\Gamma^{\kappa}_{\mu\nu}$  determining the autoparallel behavior of free particles can be split into:

The trace-free projective parameters  $\Pi^{\kappa}_{\mu\nu}$ , which determine the autoparallel paths.

The trace of the affine connection  $\Theta_{\mu} = \Gamma^{\nu}_{\mu\nu}$ , which determines a parameterization, turning the paths into autoparallel curves

#### **Non-Flat Affine Connection**

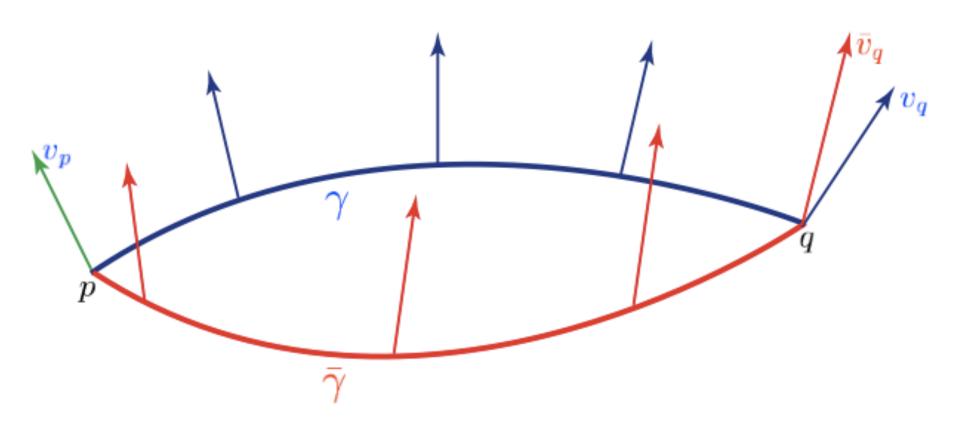
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We assume the connection is symmetric:

The torsion tensor 
$$S^{\kappa}_{[\mu\nu]} = \Gamma^{\kappa}_{[\mu\nu]} = 0$$
.

# **Affine Curvature**



#### **Affine Curvature Tensor**

The affine curvature tensor is

$$A_{\nu\mu\lambda}{}^{\kappa} \equiv 2\partial_{[\nu}\Gamma^{\kappa}{}_{\mu]\lambda} + 2\Gamma^{\kappa}{}_{[\nu|\rho|}\Gamma^{\rho}{}_{\mu]\lambda}$$

The affine Ricci tensor

$$A_{\mu\lambda} \equiv A_{\kappa\mu\lambda}{}^{\kappa}$$
,

The homothetic curvature tensor

$$V_{[\nu\mu]} \equiv A_{\nu\mu\kappa}^{\quad \kappa}$$
.

 $A_{u\lambda}$  is not necessarily symmetric, indeed:

$$A_{[\mu\lambda]} = -1/2V_{[\mu\lambda]}$$
.

# **Curvature and the Inertio- Gravitational Field**

Does the vanishing of the curvature tensor imply the vanishing of the inertio-gravitational field? Many current authors say "Yes" **Einstein** said a resounding: "No!!"

### Einstein to von Laue 1950

"It is true that in that case [the components of the curvature tensor] vanish, so one might say: "There is no gravitational field present." However, what characterizes the existence of a gravitational field from the empirical standpoint is the non-vanishing of the  $\Gamma'_{ik}$  ...

#### Einstein to von Laue 1950

If one does not think in such intuitive (anschaulich) ways, one cannot comprehend (begreifen) why something like curvature should have anything to do with gravitation in the first place. In that case, no reasonable person would have hit upon anything. The key to understanding the equality of inertial and gravitational mass would have been missing."

### The Volume Structure V

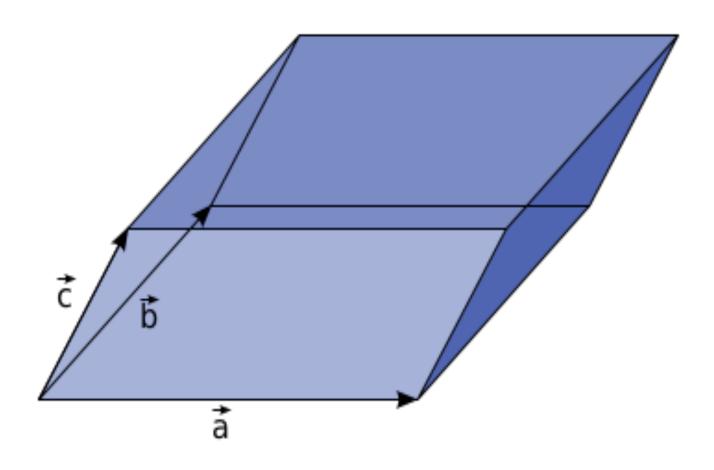
Picks out an equivalence class of sets of basis vectors defining a unit four-volume in space-time.

Two sets are equivalent if they are related by an equiaffine

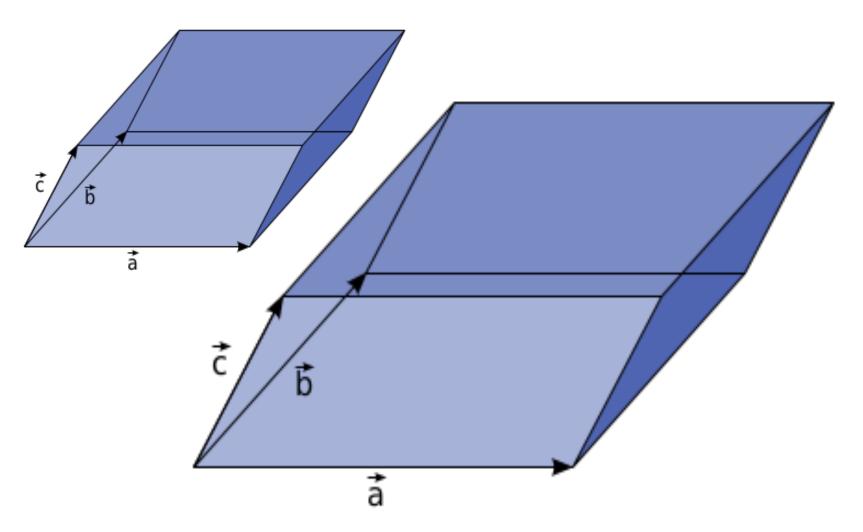
transformation  $\varepsilon SL(4, R)$ 

## The Volume Structure V

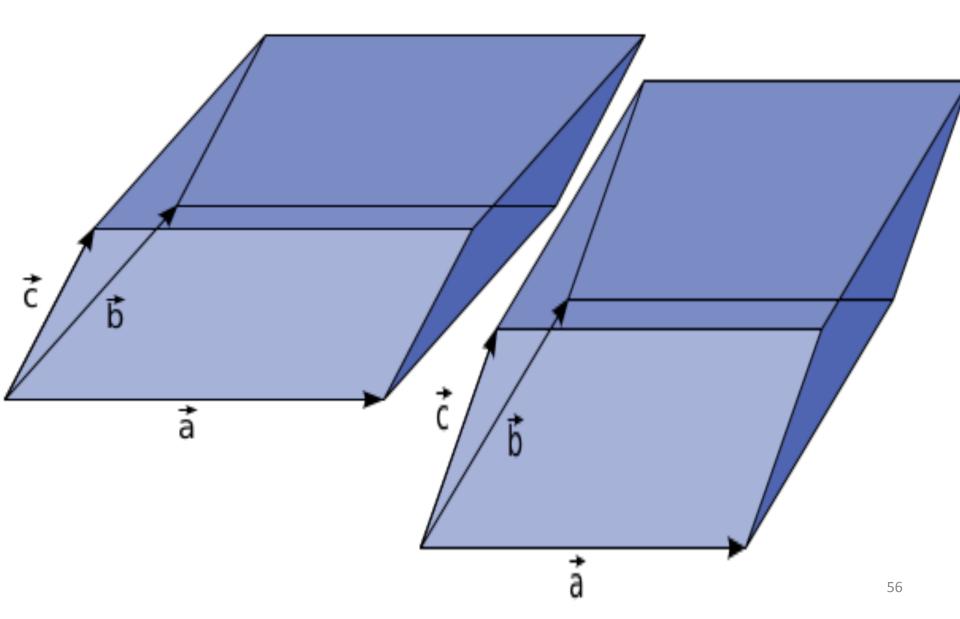




# **Similarity Transformation**



# **Equiaffine Transformation**



# Compatibility Conditions: Equi-Affine Spaces

If the affine connection and the first prolongation of the volume structure are compatible, then parallel transport of volumes is path-independent, and the connection is called equiaffine, with symmetry group SAff(4).

# Compatibility Conditions: Metric and Connection

GR includes compatibility conditions between the chrono-geometry and the inertio-gravitational field.

The usual mathematical form:

The covariant derivative of the covariant metric with respect to the connection vanishes:

$$D_{\kappa}g_{\mu\nu}=0.$$

# **Compatibility Condition**

$$D_{\kappa}g_{\mu\nu}=0$$

The usual physical interpretation:

- 1) proper spatial and temporal intervals are preferred parameters along affine autoparallel (now also geodesic) curves;
- 2) rods and clocks still read proper spatial and temporal intervals as they move in the inertio-gravitational field.

# Form of Compatibility Condition From the Lagrangian

Let 
$$g^{\mu\nu} = (-g^{\#})^{-1/2}g^{\mu\nu}$$
. The Lagrangian density is: 
$$\mathcal{L} = g^{\mu\nu}A_{\mu\nu}(\Gamma, \partial\Gamma).$$

Variation of the torsion-free connection  $\Gamma_{\mu\lambda}^{\nu}$ :

$$\delta A_{\mu\lambda} = D_{\nu} (\delta \Gamma^{\nu}_{\mu\lambda}), \text{ so}$$

$$\delta \mathcal{L} = g^{\mu\nu} \delta A_{\mu\nu}$$

$$= D_{\kappa} [g^{\mu\nu} \delta \Gamma^{\kappa}_{\mu\nu}] - D_{\kappa} (g^{\mu\nu}) \delta \Gamma^{\kappa}_{\mu\nu}.$$

The first term is a total divergence and can be eliminated. So the resulting equation is:

# **Compatibility Condition**

The form of the compatibility condition is

$$D_{\kappa}g^{\mu\nu}=0,$$

a condition on the contravariant metric density:

$$= \partial_{\kappa} g^{\mu\nu} + \Gamma^{\nu}_{\kappa\lambda} g^{\mu\lambda} + \Gamma^{\mu}_{\kappa\lambda} g^{\lambda\nu} - \Gamma_{\kappa} g^{\mu\nu}.$$

Contracting with the inverse  $g_{\mu\nu}$ ,

$$\Gamma_{\kappa} = 1/2 g_{\mu\nu} \partial_{\kappa} g^{\mu\nu} = \partial_{\kappa} \ln \left(-g^{\#}\right)^{-1/2}$$

which is the metric equiaffine condition.

# Unimodular Transformations

An equiaffine transformation in the tangent space  $T_xM$  corresponds to a unimodular transformation in the space-time manifold M. The group SL(4, R) of equiaffine transformations corresponds to the group SDIff(M) of unimodular transformations, which also have determinant = 1.

# Invariance Under *What* Group of Diffeomorphisms?

"It is to be emphasized that we have no sort of justification for general covariance of the gravitational equations. ... [W]e do not know if there is a general group of transformations, with respect to which the equations are covariant ... The question of the existence of such a group ... is the most important one ..." (Einstein 1913)

1915

XLIV

#### SITZUNGSBERICHTE

DER

KÖNIGLICH PREUSSISCHEN

#### AKADEMIE DER WISSENSCHAFTEN

Gesamtsitzung am 4. November. (S. 777) Einstein: Zur allgemeinen Relativitätstheorie. (S. 778)

#### BERLIN 1915

VERLAG DER KÖNIGLICHEN AKADEMIE DER WISSENSCHAFTEN

IN KOMMISSION BEI GEORG REIMER

# Einstein hadit right on4 November1915

# **Einstein Adopts Volume- Preserving Transformations**

"Just as the special theory of relativity is based upon the postulate that all equations have to be covariant relative to linear orthogonal transformations, so the theory developed here rests upon the postulate of the covariance of all systems of equations relative to transformations with the substitution determinant 1."

### On General Relativity Theory

"Because of the scalar character of  $\sqrt{-g}$ , one can simplify the basic formulas of the formation of covariants, as compared to those of general covariance; which in short means, the factors  $\sqrt{-g}$ and  $1/\sqrt{-g}$  no longer occur in the basic formulas, and the distinction between tensors and V-tensors drops out."

#### **Unimodular Transformations**

Today, we call "transformations with the substitution determinant 1" unimodular transformations, and there are strong arguments for their adoption on the basis of numerous theories, both mathematical and physical, including all theories of gravitation based on the equivalence principle.

# The Big Question

Why did Einstein move back from unimodular transformations to arbitrary diffeomorphisms?

No time to discuss this question here. See papers by JS and Kaća Bradonjić

# **UCPR** Lagrangian

The Lagrangian density is now:

$$\mathcal{L} = e^{\varphi c} g^{\mu\nu} A_{\mu\nu} (\Pi^{\kappa}_{\mu\nu}, \Theta_{\mu}),$$

and variations are to be taken w.r.t.

```
e^{\varphi}(1), ^{c}g^{\mu\nu}(9) field equations,
```

 $\Pi^{\kappa}_{\mu\nu}$  (36) compatibility conditions

 $\Theta_{\mu}$  (4) equiaffine condition.

# **Unimodular Group**

Under the unimodular group,  $g^{\mu\nu}$  splits into a scalar field  $e^{-\varphi}$  and a conformal metric  ${}^cg^{\mu\nu}$ :

$$\mathbf{g}^{\mu\nu} = e^{-\boldsymbol{\varphi}} {}^{\boldsymbol{c}} \mathbf{g}^{\mu\nu}$$
.

 ${}^{c}g^{\mu\nu}$  is a tensor with det = -1, and now represents the physical metric.

So: 
$$-g^{\#} = e^{-4\varphi}$$
, In  $(-g^{\#})^{-1/2} = 2\varphi$  and  $\Gamma_{\kappa} = 2\partial_{\kappa}\varphi$ 

is the equiaffine condition.

# **Unimodular Group**

The connection  $\Gamma^{\mu}_{\nu\kappa}$  splits into a trace-free part, which is now the projective connection  $\Pi^{\mu}_{\nu\kappa}$  and its trace, which is now a one form (covariant vector)  $\Theta_{\kappa}$ .

The equiaffine condition is now

$$\Theta_{\kappa} = 2\partial_{\kappa} \varphi$$

which is quite independent of the conformal and projective structures.

# **Under Unimodular Transformations:**

The conformal metric  ${}^cg^{\mu\nu}$  determines Null hypersurfaces; together with the scalar field  $e^{\varphi}$  it governs the: propagation of massless fields.

The projective connection  $\Pi^{\mu}_{\nu\kappa}$  determines Autoparallel paths; together with the one form  $\Theta_{\kappa}$  it governs the:

behavior of massive bodies.

## Superpotentials

There are a denumerably infinite number of superpotentials, tensor densities differing in weight, from which the Einstein field equations may be deduced. Only one is conformally invariant, the one that defines a gravitational angular momentum complex such that the total angular momentum transforms as a free antisymmetric tensor.

# **Gravitational Superpotential in UCPR**

In UCPR this superpotential:

$$S^{[\beta\kappa][\lambda\alpha]} = ({}^{c}g^{\beta\alpha} {}^{c}g^{\kappa\lambda} - {}^{c}g^{\beta\lambda} {}^{c}g^{\kappa\alpha}),$$

which had been a tensor density, is the one formed naturally from the contravariant conformal metric  ${}^cg^{\mu\nu}$  and hence is a tensor—another argument for UCPR!

# Some Things Einstein Got Right!! (but only later one)

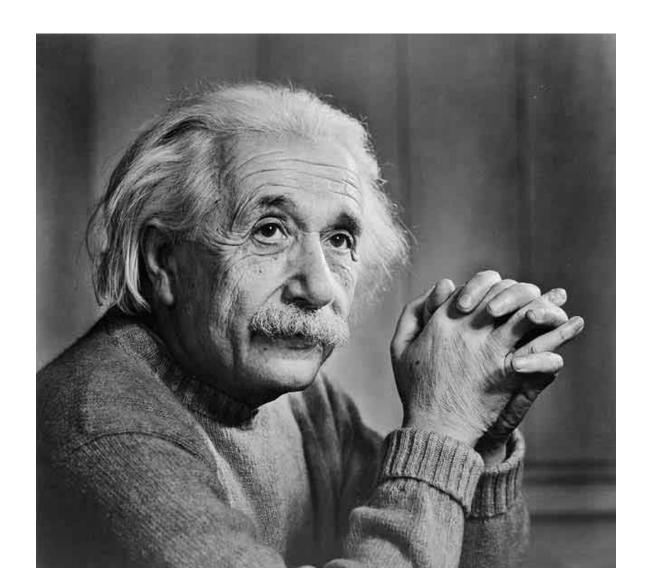
- 2) Space-Time Structures
- a) chrono-geometry— first historically; it is represented mathematically by a metric tensor.

Equivalence principle: Inertia and Gravitation are "wesensgleich" – essentially the same.

b) inertio-gravitational field—first logically; it is represented mathematically by a linear connection.

## Don't take my word-

Ask the maestro



## Einstein, *The Meaning of Relativity,* Appendix to 5<sup>th</sup> ed (1955)

The development ... of the mathematical theories essential for the setting up of general relativity had the result that at first the Riemannian metric [chronogeometry]was considered the fundamental concept on which the general theory of relativity and thus the avoidance of the inertial system were based.

## Einstein, *The Meaning of Relativity,* Appendix to 5<sup>th</sup> ed (1955)

Later, however, Levi-Civita rightly pointed out that the element of the theory that makes it possible to avoid the inertial system is rather the infinitesimal displacement field  $\Gamma^{\prime}_{ik}$ [the inertio-gravitational field]. The metric or the symmetric tensor field  $g_{ik}$  which defines it is only indirectly connected with the avoidance of the inertial system in so far as it determines a displacement field.

#### Einstein's last publication, April 1955

[I]t seems to me that Levi-Civita's most important contribution lies in the following theoretical discovery: the most essential theoretical accomplishment of general relativity, namely the elimination of "rigid" space, i.e. of the inertial system, is only *indirectly* connected with the introduction of a Riemannian metric.

#### Einstein's last publication, April 1955

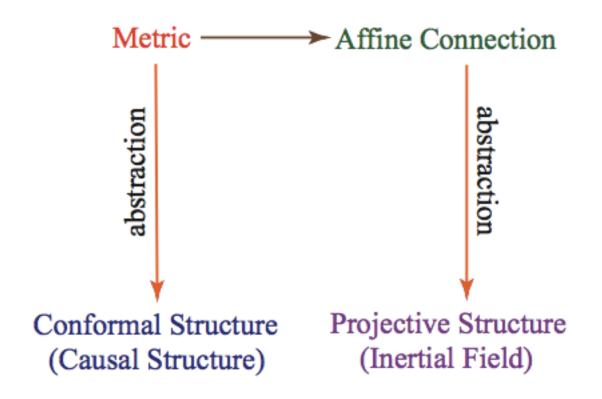
The immediately essential conceptual element is the "displacement field" ( $\Gamma_{ik}^l$ ), which expresses the infinitesimal displacement of vectors. More explicitly, it replaces the parallelism of spatially separated vectors, posited with the help of an inertial system, by an infinitesimal operation. ... In contrast to this, in a certain sense it is of secondary importance that a particular **I**-field may be derived from the existence of a Riemannian metric.

### **Space-Time Structures**

Any theory of gravitation— Newtonian, SR, GR— incorporating the equivalence principle must start from an inertio-gravitational field

These theories can only differ in the relation between this inertio-gravitational field and the chrono-geometry

## As We Have Seen, They Are Not Irreducible



#### **Curvature Tensors in UCPR**

1) The near field, tied to the sources of the inertio-gravitational field, is best described by

the projective curvature tensor.

2) The far field, needed to discuss gravitational radiation that has escaped from its sources, is best described by the conformal curvature tensor.

#### **Curvature Tensors in UCPR**

In contrast to the traditional ones, the conformal and projective curvature tensors defined in UCPR have non-vanishing Ricci tensors; so they can play a direct role in the field equations.

## **Tangent Space**

SR holds in the tangent plane, its metric  $\eta$  is invariant under equiaffine transformations.

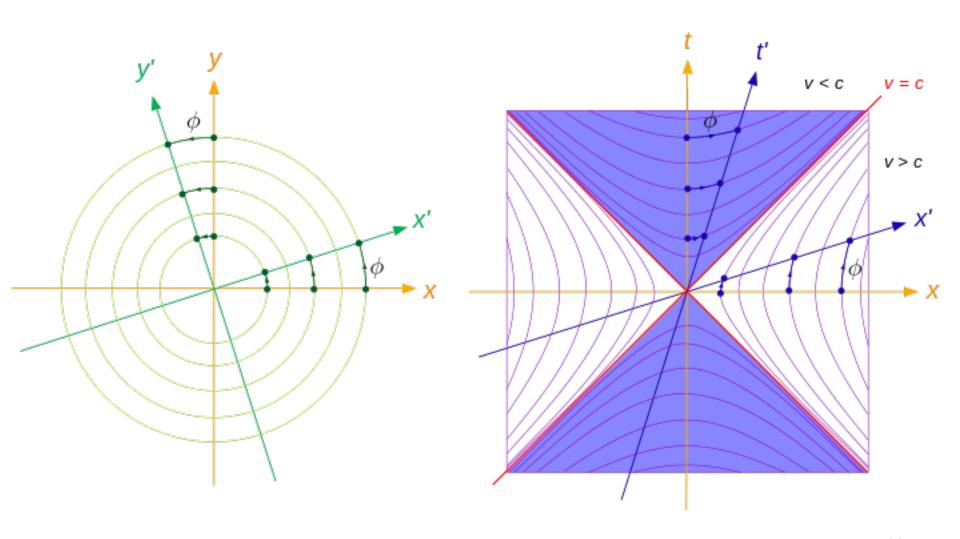
Two basic types of tetrads:

On a time-like two-plane with units:

one time-like  $e_t$ , one space-like  $e_s$ two null  $k = 1/\sqrt{2(e_t + e_s)}$ ,  $l = 1/\sqrt{2(e_t - e_s)}$ 

Plus two orthogonal space-like vectors on the orthogonal space-like two plane

### Time-Like Two-Plane

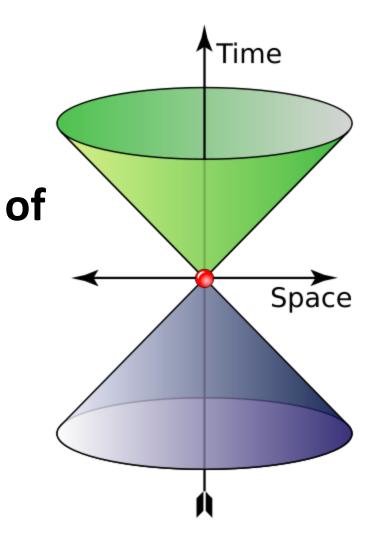


#### **Conformal Structure**

Light cone Angles

between vectors same type

Ratios of "lengths"



#### **Conformal Structure**

If one abstracts from the scale-changing property of the metric, one gets the conformal structure on the manifold.

Physically, this conformal structure represents the causal structure of spacetime. The conformal structure of spacetime determines the phases of massless test fields.

#### **Under Unimodular Transformations:**

- The (conformal) metric  ${}^{c}g^{\mu\nu}$  determines Null hypersurfaces,
- Plus the scalar field  $e^{\varphi}$  it governs the propagation of massless fields
- The (projective) connection  $\Pi^{\mu}_{\nu\kappa}$  determines Autoparallel paths,
- Plus the one form  $\Theta_{\kappa}$  it governs the behavior of massive bodies.

# The Eikonal Equation for Light

$$^{c}g^{\mu\nu}\Phi,_{\mu}\Phi,_{\nu}=0$$
,

where  $\phi$  is the phase of the light wave, is invariant under conformal transformations

Light rays are the bicharacteristics of this equation

#### **Characteristic Hypersurfaces**

The characteristic hypersurfaces  $\phi = const$  are null since:

$${}^{c}g^{\mu\nu}\Phi_{,\mu}\Phi_{,\nu}={}^{c}g^{\mu\nu}k_{\mu}k_{\nu}=0,$$

in which  $k_{\mu} = \partial_{\mu} \Phi$  is the null covector

"normal" to the hypersurface. But

$$ck^{\mu} = cg^{\mu\nu}k_{\nu}$$

actually lies on the null hypersurface:

$$^{c}k^{\mu}\Phi$$
,  $_{\mu}=0$ .

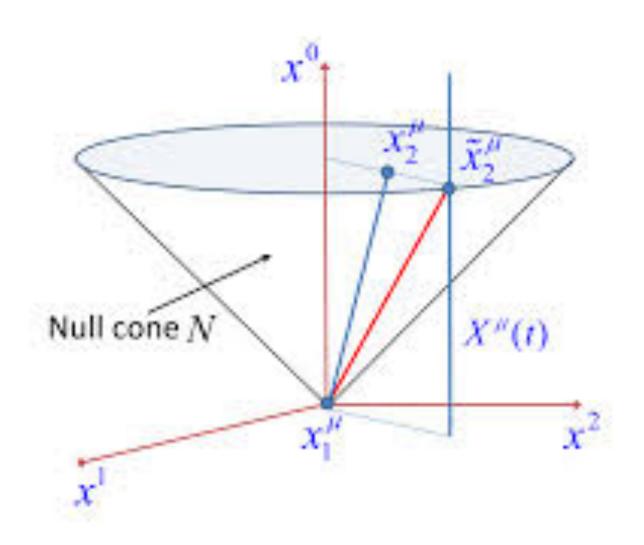
## **Bicharacteristic Rays**

The conformal metric does not determine the bicharacteristics: We may multiply  ${}^{c}k^{\mu}$  by any non-vanishing scalar function  $e^{\varphi}$  and define

$$k^{\mu}=e^{\varphi c}k^{\mu}$$
.

To define periods and wavelengths, we need to introduce the scalar volume function  $e^{\varphi}$  and define the metric  $g^{\mu\nu}=e^{\varphi} {}^{c}g^{\mu\nu}$ .

# Temporal and Spatial Projections of a Null Vector



## **Symplectic Geometry**

At this level of *symplectic geometry*, wave fronts have Phases, and rays have periods (frequencies) and wave lengths relative to some frame of reference.

Given an orthonormal basis in this f.o.r., we project  $k^{\mu}$  onto the time-like vector to get the period T:

$$k^{\mu}e^{(t)}_{\mu}=cT$$

onto a space-like vector to get the wave length  $\lambda$ :

$$k^{\mu}e^{(s)}_{\mu}=\lambda.$$

Since  $k^{\mu}$  is the diagonal of a square, the two must be numerically equal:  $cT = \lambda$ . And since  $T = 1/\nu$ ,  $\lambda \nu = c$ .

# Highest Level of Abstraction: Contact geometry

One can formulate:

Huygen's principle,

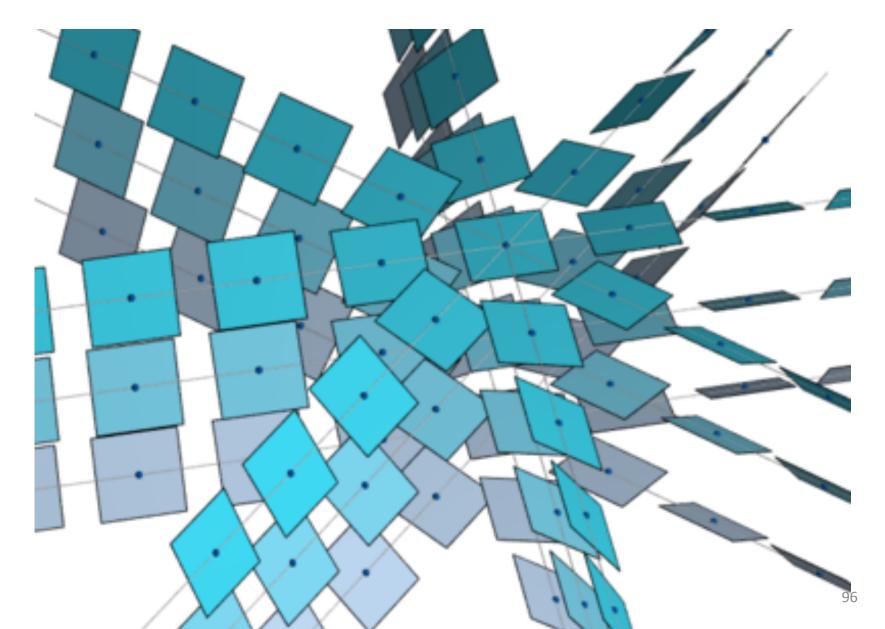
characteristic wave fronts,



Fermat's principle,

bicharacteristic rays

## **Contact Structure**

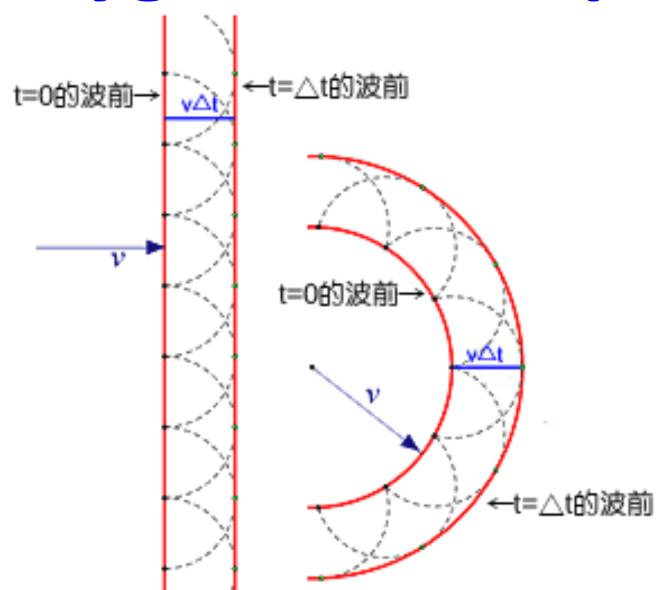


## Mechanics vs Optics

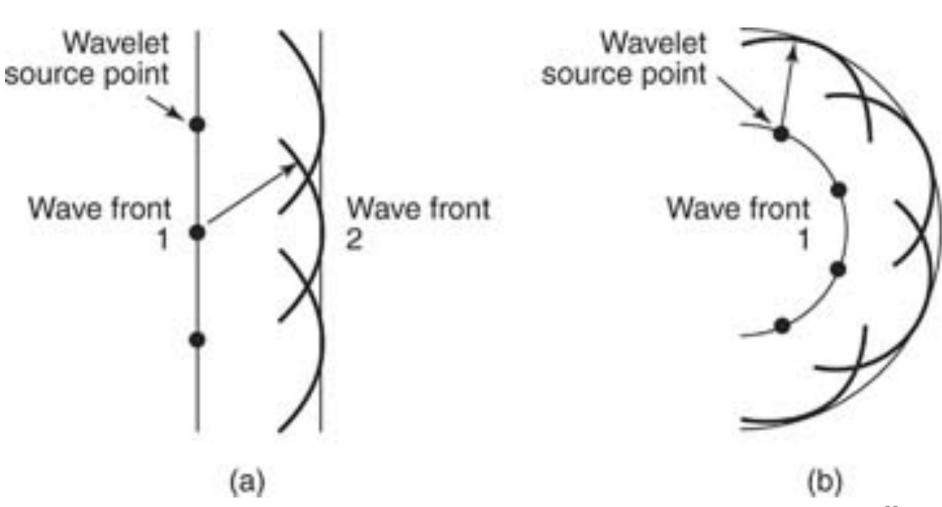
Mechanics: Spacelike Hypersurface— the normal leads from one hypersurface to the next

Optics: Null Hypersurface— the normal lies on the hypersurface, a ray leads from one wave front to the next

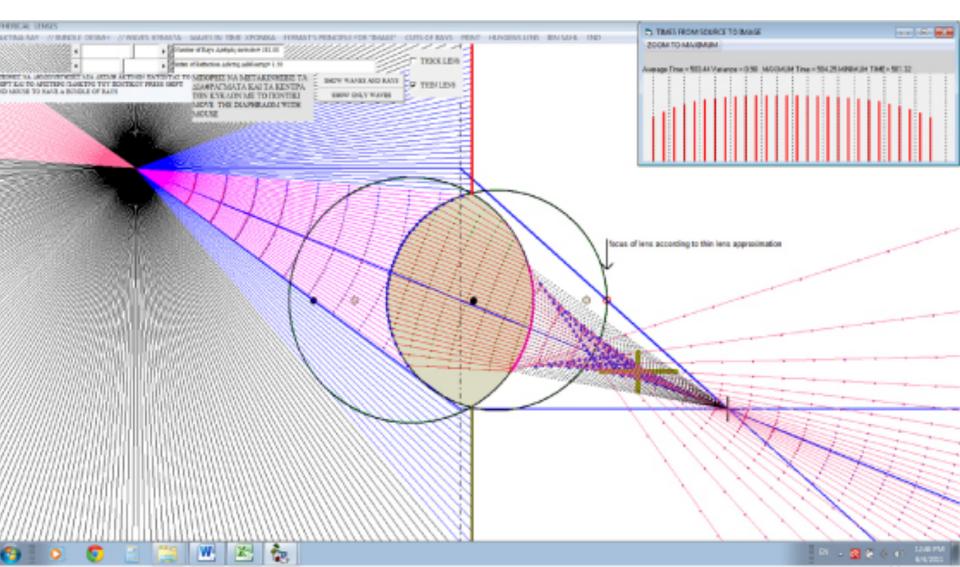
## Huygens' Principle



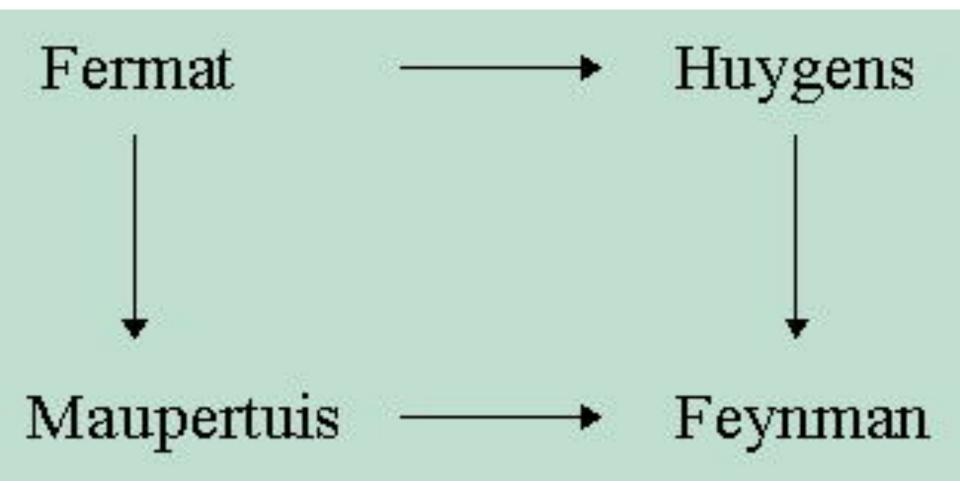
## Huygens' Principle



#### **Huygen's & Fermat's Principles**

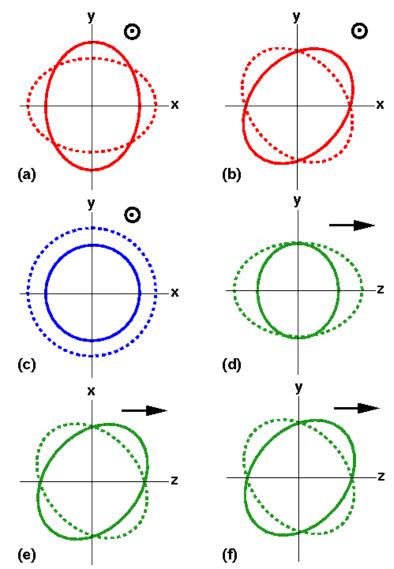


# Relations Between The Four Principles



### **Far Field-Pure Radiation**

#### **Gravitational-Wave Polarization**



#### **Far Field- Pure Radiation**

Conformal Metric: Conformal curvature tensor

Weak Field Approx'n: Assume the gravitational field is weak, so we can treat it as a perturbation around a flat Minkowski space-time metric

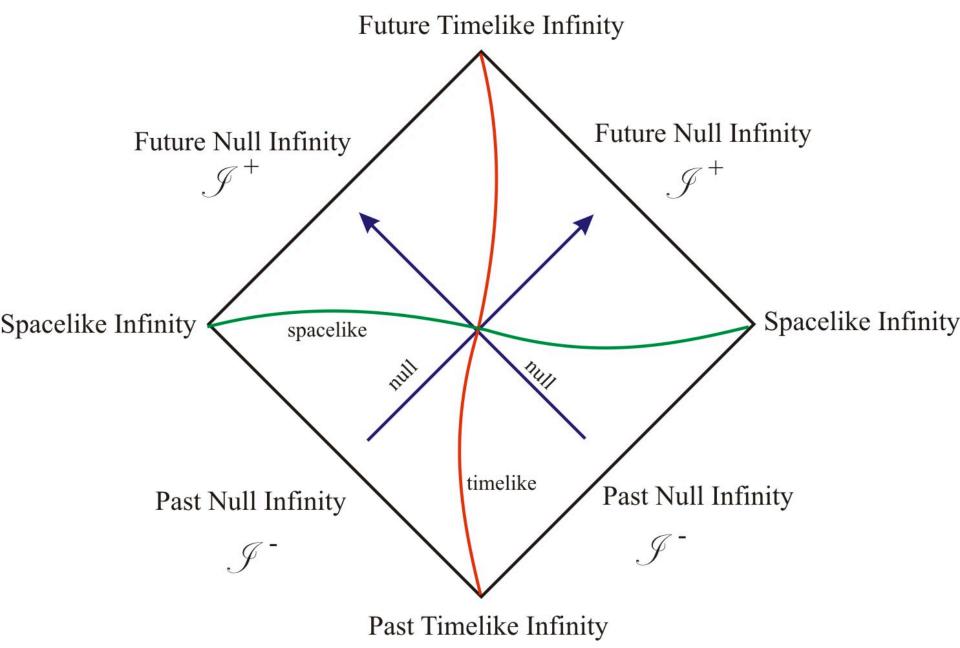
### **Asymptotic Quantization**

The Bondi-Metzner-Sachs field of any bounded radiating source asymptotically approaches the "free" gravitational radiation field, which is represented by the conformal structure at null infinity as defined by Penrose. This field has been quantized by techniques anticipated by Komar and developed in full detail by Ashtekar. The resulting asymptotic gravitons are representations of the Bondi-Metzner-Sachs group.

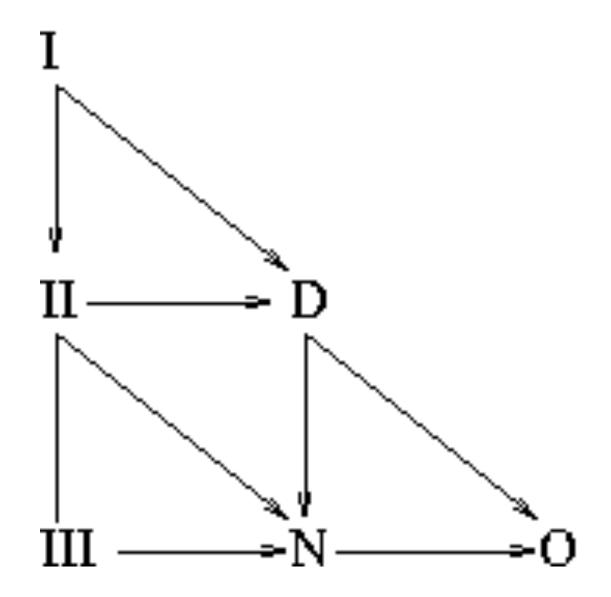
### Abhay Ashtekar: Asymptotic Quantization



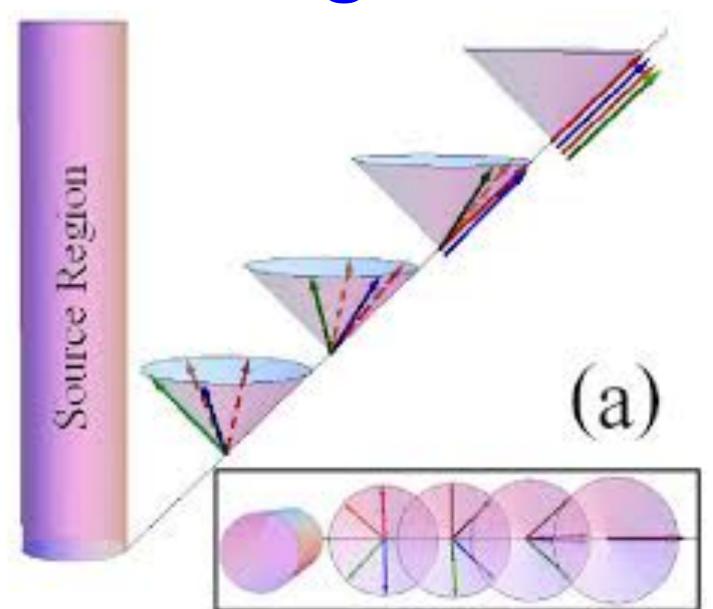




### **Petrov-Pirani Classification**



## The Peeling Theorem



#### **Asymptotic Quantization**

One can treat "pure" radiation fields independently of their sources; one needs only conformal structures to formulate the results mathematically, interpret them physically, and describe procedures for their measurement. it should thus be possible to quantize these fields directly. The homogeneous Maxwell equations are conformally invariant, so one should also be able to treat interacting "free" electromagnetic and gravitational fields by conformal techniques.

#### Near Field—Sources

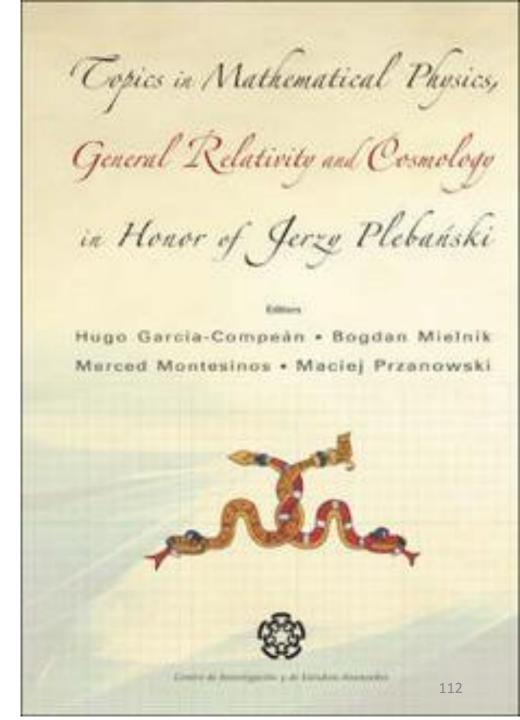
Projective Connection: Projective curvature tensor

Post-Newtonian Approx'n:
assume motions of sources are
slow, so that we can treat the
gravitational field as a perturbation about a Newtonian solution.

## Newtonian Theory-Modern Version

Even at the Newtonian level, gravitation is not an external force acting on bodies, but a modification of the hitherto fixed inertial structure of space-time, which now becomes a dynamical structure, the inertiogravitational field.

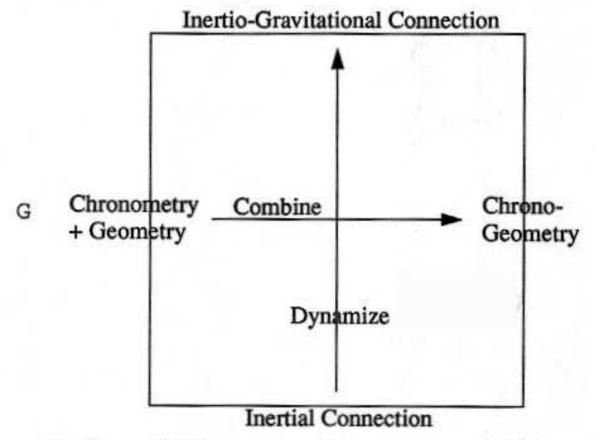
"Einstein's Intuition & the **Post-Newtonian** Approximation" (JS, published 2006, based on paper given in 2002)



#### The Bronstein Square

Gen. Non-Rel. S-T (G)

Gen. Rel. S-T (c,G)



Galilean S-T

1

Minkowski S-T (c)

[A]Ithough ... it leads us beyond Newton's original theory, the following definition seems to me to do no violence to the concept of a Newtonian-style gravitational theory: We shall require Newtonian chronometry and geometry and the compatibility conditions between both and the inertio-gravitational connection to hold. In other words, a Newtonian-style theory is one that is based on a Galileian manifold and a compatible affine connection.

Newtonian theory... allows us to go a bit further than traditional Newtonian gravitation theory. ... [A] nalysis of the compatibility conditions on the t[etrad] c[oefficients of the] c[onnection] shows that they allow the  $\Gamma^{(c)}_{(0)(0)}$  to be non-vanishing; this is well known, since they correspond physically to the electric-type Newtonian gravitational field produced by masses at rest, i.e., the  $\rho$  term in the  $T^{(0)(0)}$  component of the stress-energy tensor-- all that conventional Newtonian theory considers.

But the compatibility conditions also allow non-vanishing  $\Gamma^{(c)}_{(0)(b)}$ , which does not seem to have been noted. Physically, these components correspond to a magnetic-type Newtonian gravitational field, produced by moving masses, corresponding to the  $\rho v$  or  $T^{(0)(i)}$  components and not present in traditional Newtonian theory.

It is then simple to show that a rotating source, by creating such a magnetic-type Newtonian gravitational field  $\Gamma^{(c)}_{(0)(b)}$ , drags along the inertial frames with it. So the recent observation of the dragging of inertial frames by the rotation of the earth is a confirmation of the existence of a Newtonian magnetic-type field. We are still far from being able to observe any GR corrections to this effect.

The compatibility conditions between the dynamic inertio-gravitational connection and the fixed Euclidean geometry on the fixed absolute time hypersurfaces T = const imply that the trace of the connection vanishes.

So it is a projective connection, and it is the projective Ricci tensor that enters the gravitational field equations.

#### What Can UCPR Contribute?

Quantize the near field produced by motion of the source using the projective curvature tensor

Quantize the far radiation field using the conformal curvature tensor

Connect the two by the method of matched asymptotic expansions

#### For More Details See

"Quantum Gravity: A Heretical Vision," in S. Burra et al. (eds.), Frontiers of Fundamental Physics and Physics Education Research (Springer Proceedings in Physics 145, 2014), pp. 149-158,

and/or my Power Point:

## Quantum Gravity: A Heretical Vision

#### John Stachel

**FFP12, Udine 21-23 November 2012** 

You can contact me for comments, questions and copies at

john.stachel@gmail.com

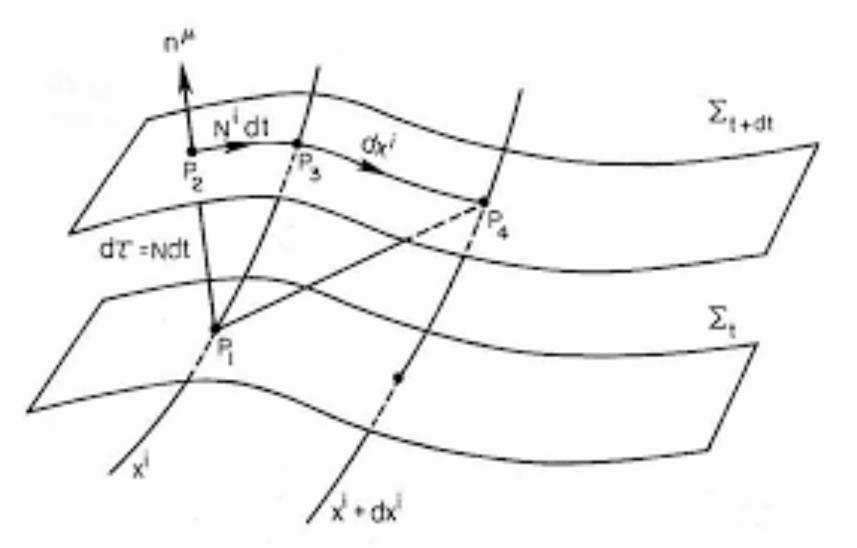
## Current Work in Progress

The role of the Frölicher-Nijenhuis Bracket in the contravariant metric approach to GR

#### **Cauchy Initial Value Problem**

Starting with the covariant metric  $g_{\mu\nu}$  and contravariant vector field  $V^{\mu}(x)$  defining a fibration of  $\mathcal{M}$ , one picks an initial hypersurface and drags it by the vector field to produce a foliation of  $\mathcal{M}$ . One then decomposes  $g_{\mu\nu}$  w.r.t. this fibration and foliation, and poses a Cauchy problem with initial data on one hypersurface: first fundamental form  $g_{\mu\nu}$  and second fundamental form  $h_{\mu\nu}$ , = -1/2 $\mathcal{L}_{V}'g_{\mu\nu}$ .

#### **Cauchy Initial Value Problem**



#### **Cauchy Initial Value Problem**

By requiring V to be a geodesic normal field n, we can show that:

$$\mathcal{L}_{n}^{2} g_{\mu\nu} = 2(h_{\mu}^{\rho} h_{\rho\nu} - B_{\mu}^{\rho} B_{\nu}^{\sigma} R_{\alpha\rho\sigma\theta} n^{\alpha} n^{\theta}).$$

 $B^{\rho}_{\mu}B^{\sigma}_{\nu}R_{\alpha\rho\sigma\theta}$   $n^{\alpha}n^{\theta}$  is determined by six field equations  $B^{\rho}_{\mu}B^{\sigma}_{\nu}R_{\rho\sigma}=0$  together with ' $g_{\mu\nu}$  and  $h_{\mu\nu}$ ; and the Gauss-Codazzi equations together with ' $g_{\mu\nu}$  and  $h_{\mu\nu}$  determine all other components of the Riemann tensor.

## Is There a *Dual* Initial Value Problem?

But suppose, starting with the contravariant metric  $g^{\mu\nu}$  and covariant vector field  $\omega_{\mu}$  (x) defining a foliation of  $\mathcal{M}$ , one picks an initial curve. Can one drag it by the covector field to produce a fibration of  $\mathcal{M}$ ?

And if so, can one then decompose  $g^{\mu\nu}$  w.r.t. the fibration and foliation, and pose an initial value problem for the fiber curve's part of  $g^{\mu\nu}$ ?

And if so what takes the place of the Lie derivative?

## Yes! There is a Dual Approach

Take three commuting vector fields  $W_{(A)}^{\mu}$  that span each of the hypersurfaces of the foliation:

$$\omega_{\mu}W_{(A)}^{\mu}=0,$$

and drag V with each of them:

$$\mathcal{L}_{W(A)}V^{\nu}$$
.

We showed this preserves the condition

$$\omega_{\mu}V^{\mu}=1.$$

#### Yes! There is a Dual Approach

Since the three  $W_{(A)}^{\mu}$  commute, the fiber created by dragging the initial fiber to any other point will be independent of the path taken between the two points.

So we have produced a fibration of  $\mathcal{M}$ . We can now show that

$$\omega_{\mu} \mathcal{L}_{W(A)} V^{\nu} = \left[\omega_{\mu} W_{(A)}, V\right]^{\nu}_{FN}.$$

Here, the Frölicher-Nijenhuis Bracket takes the place of the Lie derivative

We can show that, for a vector-valued oneform  $\omega_{\mu}W^{\nu}$ , and a vector field  $V^{\kappa}$ , the Frölicher-Nijenhuis bracket is:

$$[\boldsymbol{\omega}_{\mu} \, \boldsymbol{W}, \, \boldsymbol{V}]^{\nu}_{FN} = \boldsymbol{\omega}_{\mu} \, \mathcal{L}_{W} \boldsymbol{V}^{\nu} - (\mathcal{L}_{V} \boldsymbol{\omega}_{\mu}) \boldsymbol{W}^{\nu}.$$

If

$$\omega_{\mu}V^{\mu}=1$$
 and  $\omega_{\mu}W^{\mu}=0$ ,

this reduces to

$$[\omega_{\mu} W, V]_{FN}^{\nu} = \omega_{\mu} \mathcal{L}_{W} V^{\nu}$$

#### NATURAL OPERATIONS IN DIFFERENTIAL GEOMETRY

Ivan Kolář Peter W. Michor Jan Slovák

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Electronic relation. Originally published by Norlayer-Verlag, Berlin Beldelberg, 1960, 1960-3-540 No.20-4, 1960-0-307-56225-6.

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#### Ivan Kolář Peter W. Michor Jan Slovák

#### Natural Operations in Differential Geometry



Springer-Verlag Berlin Heidelberg GmbH

"Here we present the The Frölicher-Nijenhuis Bracket (a natural extension of the Lie bracket from vector fields to vector valued differential forms) as one of the basic structures of differential geometry and we base nearly all treatment of curvature and Bianchi identities on it."

"This allows us to present the concept of a connection first on general fiber bundles (without structure group), with curvature, parallel transport and Bianchi identity, and only then add Gequivariance as a further property for principal fiber bundles."

"We think that in this way the underlying geometric ideas are more clearly understood by the novice than in the traditional approach, where too much structure at the same time is rather confusing."

# Vielen Dank für Ihre Aufmerksamkeit